The Electron Cloud Effect: Summary

M. A. Furman LBNL Snowmass mtg. July 2001



PAC01: "smoking gun" evidence at many machines: PEP-II, KEK-B, BEPC, APS, PSR, SPS, PS, expected at LHC, possibly SNS This talk builds on the experience of many people; I am particularly grateful to: J. S. Berg, M. Blaskiewicz, O. Brüning, Y. Cai, I. Collins, A. Drees, Kulikov, G. Lambertson, R. Macek, K. Ohmi, M. Pivi, J. Rogers, R. O. Gröbner, K. Harkay, S. Heifets, N. Hilleret, R. Kirby, A. Rosenberg, J. Seeman, G. Stupakov, F. Zimmermann



The ECE in a nutshell:

- 1. beam produces electrons (PHOTOEMISSION) RESIDUAL CAS IONIZATION, PARTICLE LOSSES HITTING CHANGER WALLS
 - 2. electrons get rattled around the chamber
- from single-bunch passage, and/or
- from multi-bunch passages
- 3. number of electrons may increase or decrease from collisions against the walls (sex
- 4. electron cloud couple to the protons. BEAM
- within the bunch, and/or
- successive bunches are coupled via the interbunch electron cloud

Especially strong effect for intense, positively-charged beams (e⁺ or p); possible consequences:

- fast instability (single bunch or multibunch)
- emittance increase and/or particle losses
- excessive energy deposition on the walls (e.g., LHC)

In summary: the ECE is a consequence of the strong coupling between the beam and its environment; many ingredients: beam energy, bunch charge and spacing, photoelectric yield, photon reflectivity, secondary emission yield, chamber size and geometry, ...



BRIEF HISTORY OF E.C.E.

- 1977 : FIRST EVIDENCE; BEAM-INDUCED MULTIPACTING (BIM) O. GRÖBNER, CERN (ISR, TESTING SAMPLE ALUMINUM CHAMBERS FOR ISABELLE)
 - SUDDEN INCREASE IN VAC. PRESSURE (RESONANT ECE)
- 1995: IZAWA, SATO, TOYONASU (PHOTON FACTORY, KEK)
 - INSTABILITY FOR et BEAM BUT NOTE
 - 8 POAD SPECTRUM (NON-RESONANT ECE) (PRL 95)
- DEC 85: PSR (LANL) e-p INSTABILITY (R.MACEK) D. NEUFFER PICTURE (~ 1991)
- 2 1985 CESR "ANOMALOUS ANTI-DAMPING" (J. ROGERS EXPLANATION ~ 1996 : E TRAPPING IN FIELD OF VACUUM PUMPS)
 - SIMULATION CODES: 21995 K. OHHI; M. FURMAN 2 1997 F. ZI MMERMANN

H. BLASKIEWICZ, A.DANILOU, Y. ALEKSHUDRO

J. ROGERS

S. HEI FETS

Y. CAI

R. DAVIDSON, H. QAN T. S. WANG "2-STREAM

K. HAIRKAY & R. ROSENBERG (ANL) DEDICATED & PROBES: 2 1007

ANALYTIC: G. STUPAKOV

L. VOS

BEAM - INDUCED MULTIPACTING

(RESONANT ECE)

(0.620 MER, 1977)

FLAT CHAMBER

E TRAVERSAL TIME - At (DUE TO ONE BUNCH PASSAGE)

SET $\Delta t = \frac{s_0}{c}$

SO = BUNCH SPACING

 $\Rightarrow s_B = \frac{b^2}{N \pi_e}$

N = PARTICUES / BUNCH $\pi_e = 2.82 \times 10^{-15} \text{ m}$ $\ell = e^2 / \text{mc}^2$

(NECLECTS IMAGE FIELDS, BUNCH LENGTH, ENERGY SPREAD OF et, ...)

ECE SEEN AT

PHOTON FACTORY (PF)

BEPC:
PEP-II

KEK-B

SPS

PS

APS

PSR

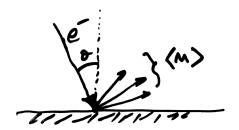
EXPECTED AT:

LHC VLHC (STAGE 2) NLC DR CLIC DR SNS (?)

Confidence level of ECE evidence (PACOI) (IMHO)

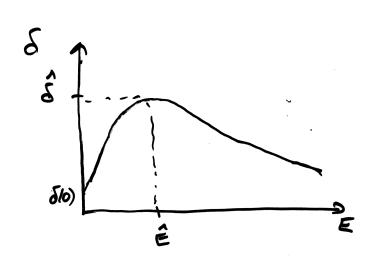
- •Wild guess
- •Guess
- •Hunch
- •Educated guess
- •Vox populi
- •The preponderance of evidence suggests that...
- •What else could it be
- •Smoking gun
- •Three unrelated witnesses in broad daylight swear that...

AT CENTER STAGE IS SEY



$$\delta(E, \theta) = \langle m \rangle$$

$$\frac{E}{8} \frac{8}{8} \frac{80}{800}$$
Tin ~250 il.1 i?
Cu ~120-250) 12-1.3! 0.6?
Alp, ~300 ?2.5 ?
SS ~250 ~2.2 ?



IN PRACTICE, WHAT COUNTS IS NOT δ BUT $\delta = \sqrt{\frac{1}{2}} dE d\Theta \delta(E, \Theta) \times O(E, \Theta)$

Contributions to the sample current from electron backscatter-bombardment of nearby surfaces (all graphite-coated) were measured at normal incidence using both RP and ZR methods. Secondaries generated at these surfaces have mostly very low energy and serve to strike and reduce the apparent yield of the sample. ZR curves need to be corrected (< 5%) for this effect. Primary electrons elastically scattering in the forward direction from the surface at grazing incidence were collected in a "black-hole" gridded structure behind the sample holder. The sample was moved laterally several electron beam diameters between successive angular measurements, to ensure that a previously-unbombarded area was used for the next SEY measurement.

XPS measurements were made on TiN-coated samples in a separate UHV system and then the samples were transferred to the SEY-measurement chamber. The SEY-measurement chamber has an Auger electron spectrometer (AES), which is used to check surface chemistry on a spot several mm away from the SEY-measurement area. Conditions in this system are similar to the LER, cleaned for UHV operation and unbaked with a pressure of one ntorr or so.

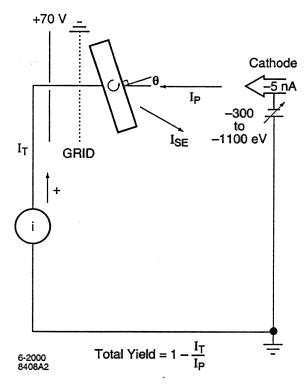
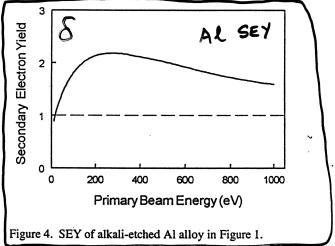


Figure 3. Circuit for zero retard yield measurements, angular beam incidence. The volume around the sample is at ground potential and field-free.

3. Results

3.1 General Comments

Figure 4 shows the unacceptably large (for beam chamber use) peak SEY of >2 of air-oxidized etched aluminum alloy. In fact, chemically removing residual carbon from the surface serves to increase the SEY by increasing the percentage of surface covered by high-yield aluminum oxide in place of lower yield carbon. Copper has a much lower yield and need not, for present machines at least, be coated with a SEY-lowering material. For aluminum alloy chambers and high-field components, TiN coating is frequently used. However, chambers are always exposed



to atmosphere following coating, which leads to a rise in the SEY of the coated surface. Such surfaces need to be "conditioned" in situ to restore a low SEY.

Conditioning of beam chambers, i.e., an in-situ reduction of the gas load from the chamber walls occurs by synchrotron light and electron bombardment. process and its effect on the SEY of the walls is characterized as follows. The as-deposited/air-exposed SEY (e.g., TiN-coated aluminum alloy, Figure 5, inset) is mainly determined by the presence of a surface water and hydrocarbon layer [9]. Electron bombardment of this layer, the initial processing that occurs during machine comissioning, desorbs the gas layer and also results in electron-induced adsorption of carbon from carboncontaining residual-gas molecules. Water desorption and carbon deposition, amongst other processes discussed later in this paper, serve to reduce the peak SEY to near one (Figure 5, graph). This yield is close to the in-situ asdeposited TiN value [4]. The conditioning time can be reduced by glow discharge cleaning (GDC) of the chamber walls after machine assembly. Such ion sputterremoval of surface gas is much more efficient than photon/electron bombardment. Measurements of the SEY following simulated conditioning and GDC are included below.

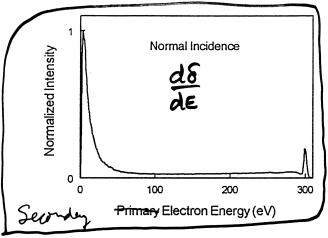


Figure 15. Collected electron energy distribution from TiN-coated Al alloy. The primary beam and analyzer axis were normal to the sample surface. The most probable secondary electron energy is 2.9 eV (8.4 eV FWHM).

The generated electron distribution curve (EDC) consists of true secondary, inelastic and elastically-reflected electrons. Forty eV is usually considered the high energy-pass cutoff for "true" secondary electrons. Relatively few primary electrons are backscattered elastically. The remainder of the generated spectrum consists of re-diffused primaries that have suffered loss events (ionizations, Augers, plasmons, etc.). The relative proportion of each kind of generated species is shown in Table I. An incidence energy of 300 eV was chosen because that is about the peak of the SEY and, hence, the most efficient for total secondary electron production. At lower primary energies, a progressively larger fraction of the distribution will be elastic [11]. The emitted electron angular spectrum will be cosine but peaked for backscattered elastics.

θ	0-40 eV	40–295 eV	295-310 eV
0°	58.9 % 56.8 %	36.6 %	4.5 %
82.5°	56.8 %	38.9 %	4.3 %

Table I. Normalized EDC intensity areas at normal (0°) and grazing (82.5°) primary electron incidence angle θ .

At a grazing incidence of 82.5°, the EDC still looks very similar to Figure 15 but now with the most probable energy at 2.7 eV (7.1 eV FWHM). Unlike normal incidence, more primaries will stay within the escape depth of generated low-energy secondaries, i.e., grazing incidence increases the SEY. Unfortunately, the energy analyzer has its axis along the primary beam direction, resulting in an off-sample-normal collection slit. The relative proportion of each kind of generated species is about the same (Table I), however, suggesting that the TiN surface roughness is sufficient to homogenize the secondary emission over all incidence angles.

3.5 The Secondary Yield Of Other Materials

The SEY of some materials used for beam vacuum chamber construction (and graphite, for reference) are plotted in Figure 16. The surfaces were Ar-sputtered (~1x10¹⁷ ions-cm⁻²) to remove most, but not all, surface contamination, as determined by Auger electron spectroscopy. Remaining carbon levels were in the 1-5 atomic per cent range. The yields of Figure 16 are those expected for beam vacuum chambers that have been glow-discharge cleaned after assembly and pumpdown.

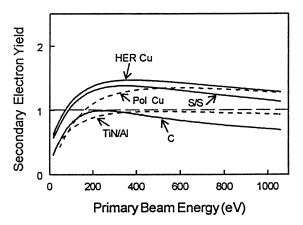


Figure 16. SEY of some materials used in accelerator/storage ring construction. Graphitic carbon is included as a measurement reference.

The yield of sputter-cleaned TiN is consistent with previous work for as-deposited, not exposed to air, films [4]. The result for polished copper ("Pol Cu") shows a shift of the peak SEY to higher primary energy, probably due to less electron-scattering (SEY-lowering) surface disorder.

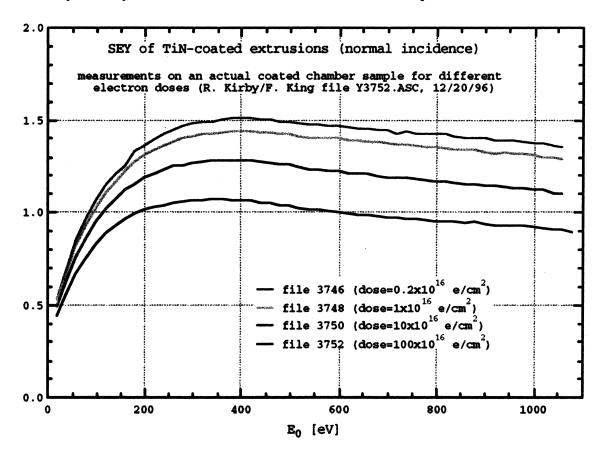
4. Discussion

4.1 Primary Angle Dependence of SEY Peak Yield

SEY theory predicts that the normalized yield δ/δ_0 , where δ_0 is the SEY maximum at normal incidence, should vary as $1/\cos\theta$ [12]. Figures 17 and 18 show that the experimental data from TiN/AI or Cu do not match theory for the simple case, $E_P = E_{MAX}$, (i.e., the yield is not tending toward large values as θ approaches grazing angles) and, incidentally, appears independent of conditioning (Figure 17) as well. E_P is the primary electron energy and E_{MAX} is the primary electron energy corresponding to the maximum SEY, δ_0 , at normal incidence.

Secondary electron emission

- main quantity of interest is the secondary emission yield (SEY) δ :
 - δ =average number of ejected electrons per incident electron
 - + function of: incident energy E_0 , incident angle θ_0 and surface material
 - model parametrized to fit data by R. Kirby and F. King (SLAC) for TiN-coated extrusion sample



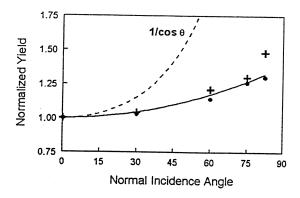


Figure 17. Normalized peak yields from TiN/Al (Figures 6 (\bullet) and 8 (+)) vs. primary beam incidence angle, at E_{MAX} =440 and 389 eV, respectively. The curve fit is exp [0.32 (1-cos θ)].

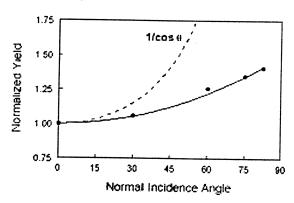


Figure 18. Normalized peak yields from HER copper (Figure 13) vs. primary beam incidence angle, at E_{MAX} =675 eV. The curve fit is exp [0.40 (1-cos θ)].

The lack of agreement with theory begins to occur far from grazing incidence, so this does not appear to be a surface-only related phenomenon. In fact, the surface finishes of the TiN/Al and Cu are quite different. However, as the primary beam angle moves toward grazing incidence, an ever-increasing fraction of primaries will exit the surface, in the forward direction, before losing their full energy to secondary electron production [13]. A simple phenomenological explanation, due to Bruining [14], accounts for the exponential dependance of Figures 17 and 18. Suppose X_m is the average depth (Figure 19) at which N_s secondary electrons are generated at normal primary beam incidence.

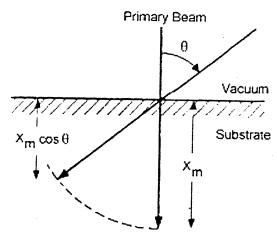


Figure 19. Schematic diagram [14] showing the secondary electron generation depth, X_m , as a function of primary electron incidence angle, θ .

Then the escape probability depends on secondary electron absorption α , and the SEY is given by

1)
$$\delta_0 = N_s e^{-\alpha X}$$

At other than normal incidence, the yield is

2)
$$\delta_{\theta} = N_s e^{-(\alpha X_m \cos \theta)}$$

Combining, the normalized yield as a function of primary incidence angle is given by

3)
$$\delta_{\theta} / \delta_0 = e^{\alpha X}_{m} (1-\cos\theta)$$

which dependence fits the experimental data very well, as shown in Figures 17 and 18.

4.2 Electron Conditioning - LER

Figure 20 shows the air-exposed TiN-coated Al peak yields, as a function of electron dose, for three different samples, all pre-exposed to ambient atmosphere. The dosing and SEY measuring beams were normal to the surface. As an aside, others [15] have made such measurements to higher dose and shown that the peak yield levels off at approximately $1x10^{18}$ electrons cm⁻².

MOTIVATION

- ELECTRON CLOUD CODES HAVE BEEN DEVELOPED FOR ~ 6 YEARS NOW.
- USED TO PREDICT ELECTRON CHOUD EFFECTS AT VARIOUS MACHINES, eg. PEP-II, LHC, VLHC, APS, PSR, SPS,...

MUNOMPHOSTER CODE

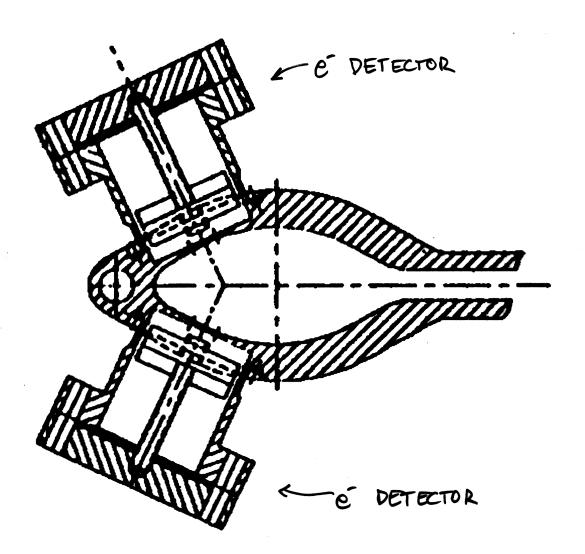
- DURING THE PAST ~4 YEARS,

 SPECIALIZED & DETECTORS HAVE BEEN
 BUILT + DEPLOYED AT APS, PSR, BER
- COMPARE SIMULATION RESULTS
 AGAINST NEASUREMENTS AT APS

SPECIFICALLY:

- * ELECTRON CURRENT AT THE WALLS
- * APS WITH et BEAMS
- * VARY BUNCH SPACING , FIXED TOTAL CURRE

APS VACUUM CHAMBER AND & PETECTORS (K. HARKAY + R. ROSENBERG)

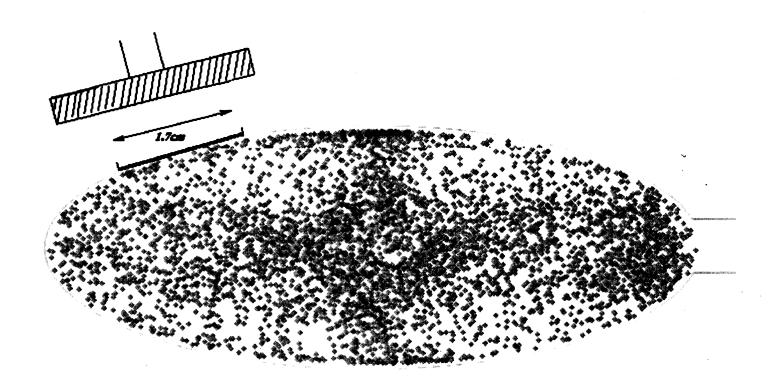


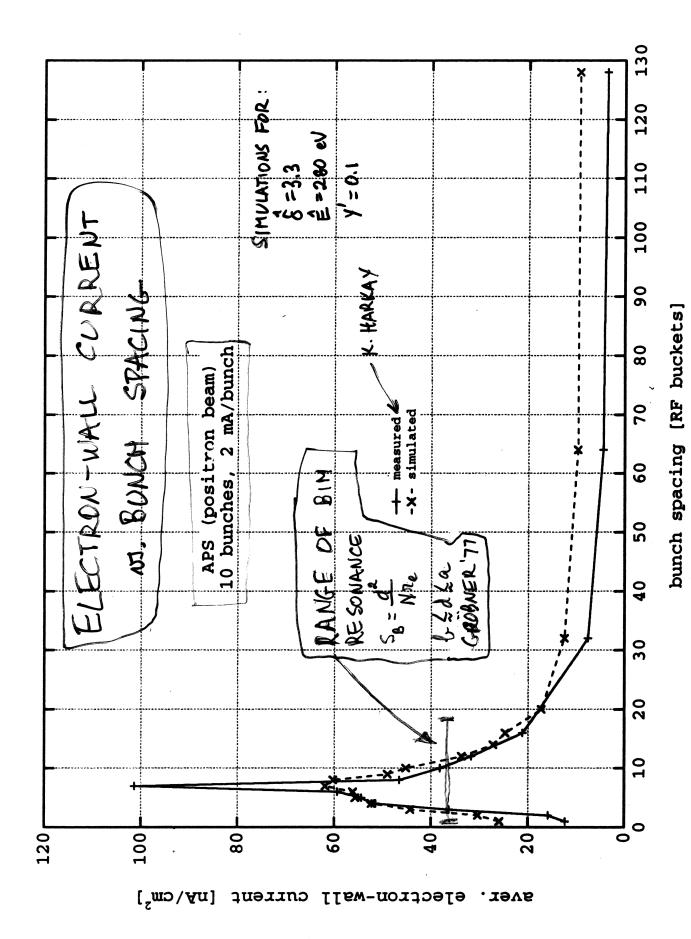
FEATURES OF THE SIMULATION CODE

- BEAM IS NON-DYNAMICAL
 - ELECTRONS IN THE CLOUD ARE FULLY DYNAMICAL
 - ELECTRONS GENERATED BY
 - * PHOTOEMISSION
 - * SECONDARY ELECTRON EMISSION
 - * RESIDUAL GAS IONIZATION
 - CAN SIMULATE:
 - * FIELD-FREE REGION
 - * DIPOLE FIELD
- 3D ELECTRON KINEMATICS
- 2D (TRANSVERSE) FORCES
- SPACE CHARGE
- IMAGE CHARGES
- ELLIPTICAL VACUUM CHAMBER
- POSSIBLE ANTE CHAMBER

δ = SECONDARY EMISSION YIELD) PETALLED do = SECONDARY ENERGY SPECTRUM) INPUT HODEL

DATA FOR SEY FROM SLAC MEASUREMENTS (R. KIRBY)





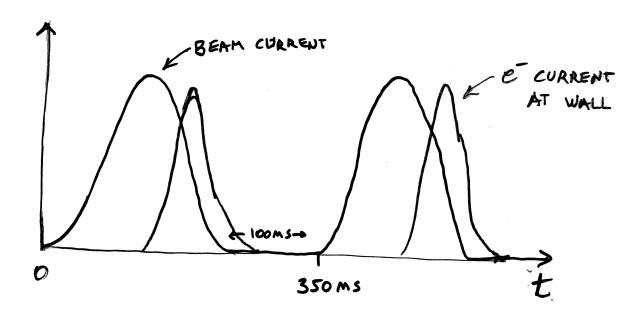
SIMULATIONS + MEASUREMENTS

AT PSR

- R+D PROGRAM TO UNDERSTAND PSR INSTABILITY

- EST. ~ FEB 2000 , FUNDED BY SNS
- LANL , BNL , LBNL , PPPL , ORNL , ANL
- P.I. R. MACER
- TERMINATED ~ DEC 2000 (FORMALLY)
- WORK CONTINUES
- POST-DOC MAURO PIVI HIRED UNDER THIS PROGRAM (LOUI,
- COMBINATION OF THEORY, SIMULATIONS 4
 EXPERIMENTS
- INSTALLED ANL & DETECTORS IN PSR (HARRAY ROSENBERG)

TRAILING-EDGE MULTIPACTING



ELECTRONS "SUCKED IN"

DURING LEADING EDGE

THEN RELEASED DURING

TRAILING EDGE

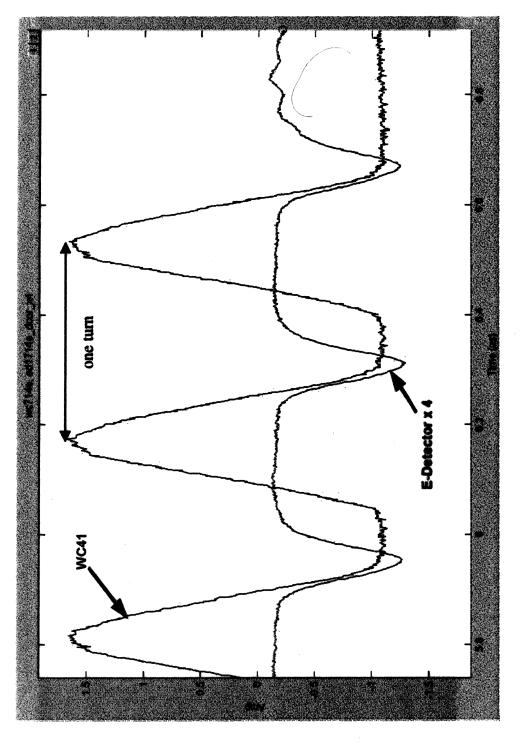
C = 90 m $N = 5 \times 10^{13}$ $T_{\text{BUNCH}} \approx 250 \text{ ms}$ $T_{\text{GAR}} \approx 100 \text{ ms}$

SIMULATIONS: BLASKIEWICZ, ALEKSANDROV, PANILOV

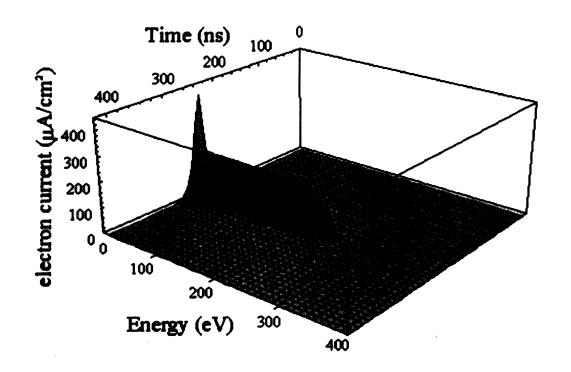
MAIN SOURCE OF E : LOST PROTONS HITTING THE CHAMBER ~ 10-6 p/p/TURN ~ 100 e/p

PSR beam current and electron-detector signal:





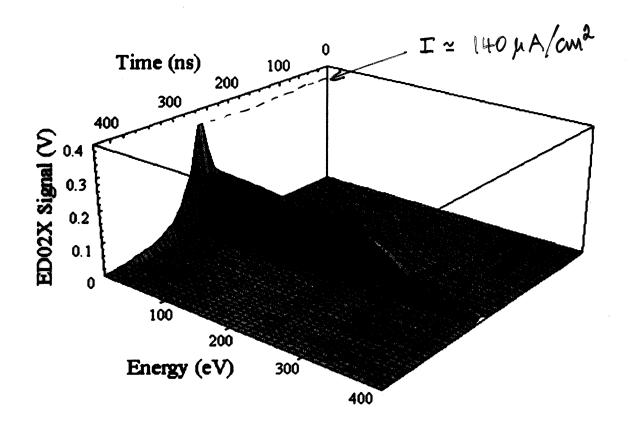




PSR: SIMULATED

CUMMULATIVE-ENERGY SPECTRUM

(M. FURMAN, M. PIVI, PACOI)



PSR! MEASURED

ROO CUMMULATIVE-ENERGY SPECTRUM

(COURTEST A. MACEK)

NEW SIMULATION RESULTS FOR THE

ELECTRON- CLOUD EFFECT AT THE

PEP-II POSITRON RING

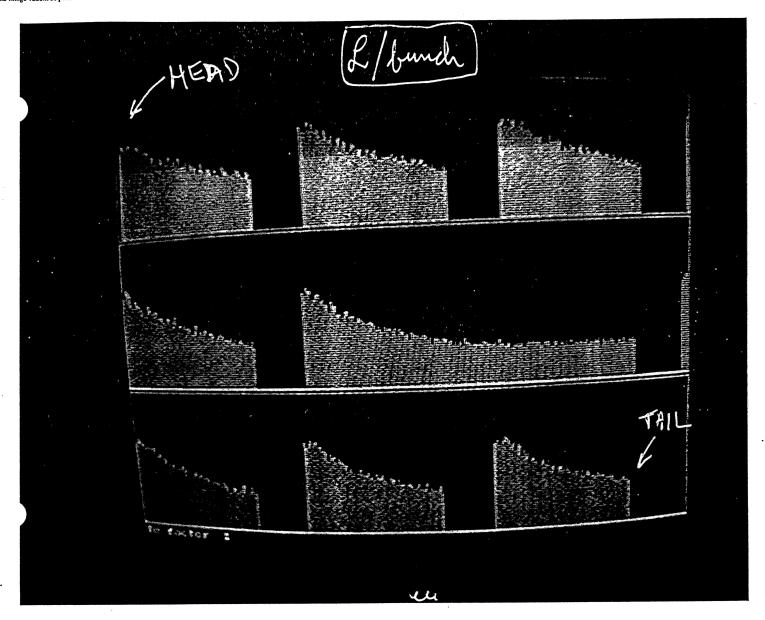
Y. CAI, S. HEIFETS, J. SEEMAN (SLAC)
M. FURMAN, M. PIVI (LBNL)

FRI. 6/22/01 PACOL

ACKNOWLEDGHENTS: R.KIRBY, N. HILLERET, I. COLLINS, V. BAGLIN, F. ZIMMERMANN, O. GRÖBNER, O. BRÜNING, G. LAMBERTSON, K. KENNEDY, K. OHMI, T. ROGERS, F. RUGGIERO, M. ZISMAN, V. BAGLIN, K. HARRAY

RELATED PRESENTATIONS AT PACOL:

TPPH126 ₹
W0AA003
TPPH100
R0PB001



This is one turn obvious not BB problem probably not RF transist PEP-IL Bunch By Bunch Luminosity

date taken July 2000 (most diluly)

(J. SEEMAN)

GRADUAL & DECREASE ALONG BUNCH TRAIN
IS DUE TO et BEAM BLOWUP DUE TO
ELECTRON CLOUD

STRONG EVIDENCE FROM:

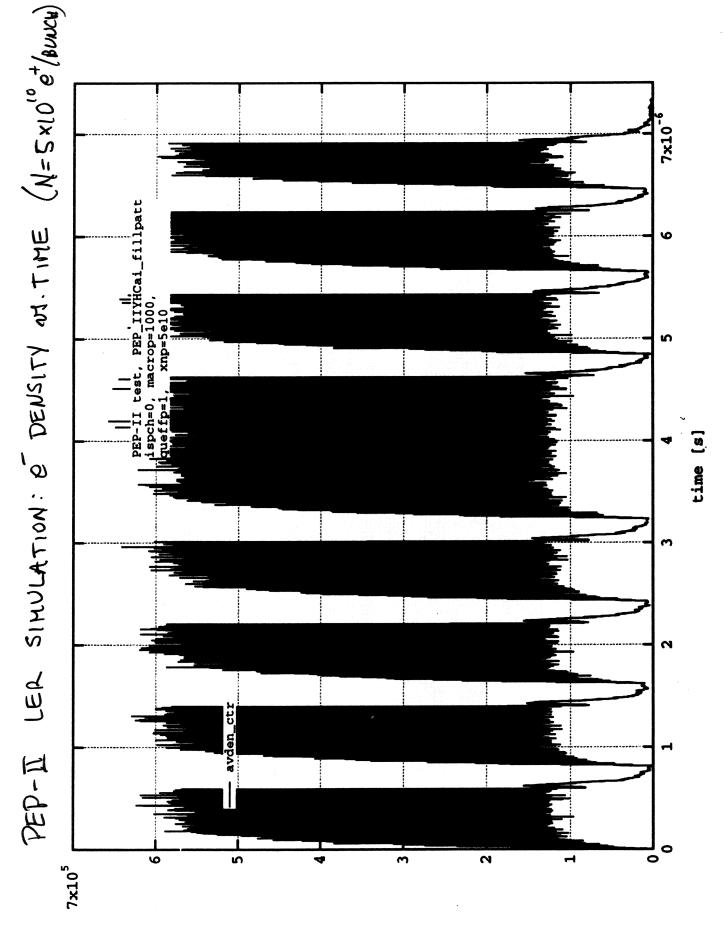
- PRESSURE RISE NO. I (FILL PATT. DEPENDE
- BENEFICIAL EFFECT OF SOLENOIDAL FIELD
- 0x + 0y BLOWUP NO. I (SYNCHROTRON LIGHT INTERFEROMETER)

SIMULATIONS WERE DONE BY TWO CODES:

- 1) LBNL CODE -> DETERMINE ELECTRON-CLOUD DENSITY Me
- 2) SLAC CODE: INPUT Me -> DETERMINE
 BEAM BLOWUP

* FEATURES OF THE LBNL CODE:

- ARBITRARY FILL PATTERN
- RIGID BUNCHES
- PHOTO ELECTRONS
- SECONDARY ELECTRON EMISSION
- VAC. CHAMBER GEONETRY, INCL. ANTECHAMBER
- 3D ELECTRON KINEMATICS
- 2D ETRANSVERSE) FORCES
- SPACE CHARGE
- I MAGE CHARGES



ejectrons/cm**3

FEATURES OF THE SLAC CODE - MODEL: OHMI 4 ZIMMERMANN (PRL 85,3821,2000)

- ELECTRON CLOUD IS STATIC, LUMPED AT ONE POINT IN THE RING

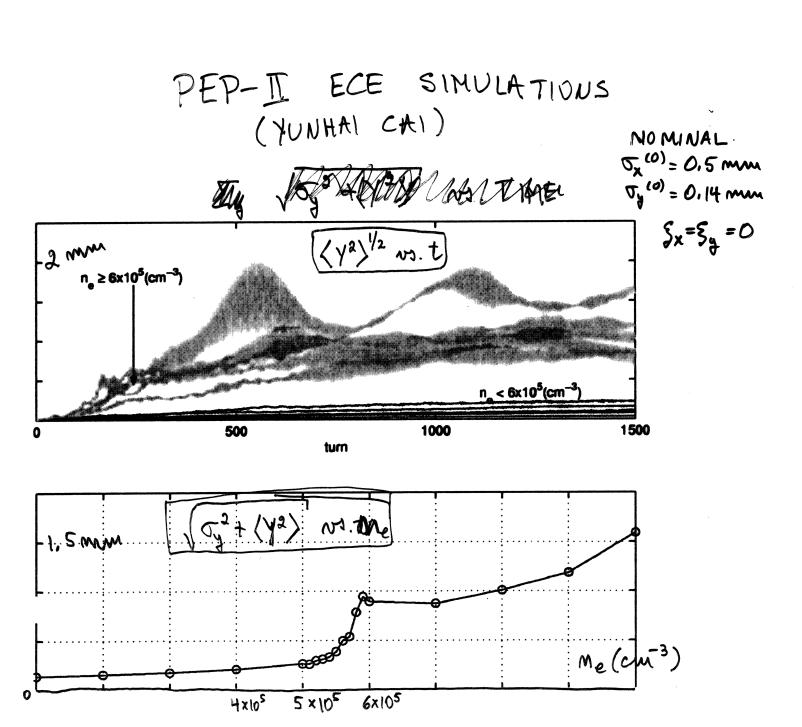
- POSITRON BUNCH IS DYNAMICAL (MACROPARTICLES, LONGITUDINALLY SLICED)

- CHROHATICITY

- 3D (INCL. SYNCHROTRON MOTION OF POSITRONS)

- C++

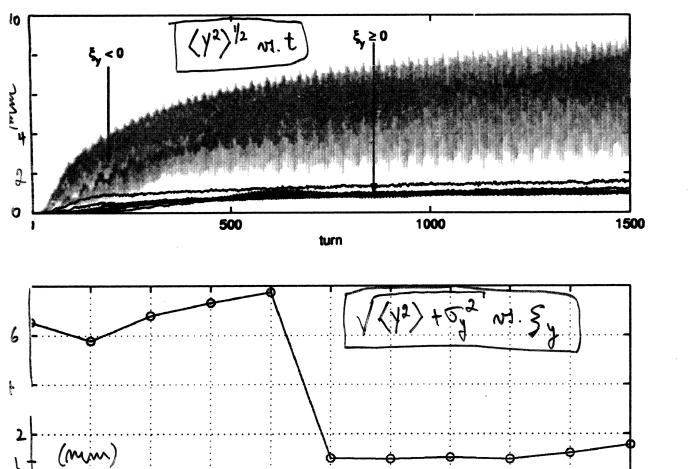
		~
E	BEAM EVERGY	3.1 GeV
Bx	AV. Bx	9.37 M
Ba	AV. By	12.47 M
Q ^X	AV. ox	0.5 mm
	N. 5y	0.14 mm
5	RMS BUNCH LENGTH	1.3 cm
Jp/p	MOM. SPREAD	7.7×10-4
X	H. TUNE	0.649
7/4	J. TUNE	0.564
75	SYNCHR. TUNE	0.0251
Np	# et /BUNCH	1×10"
Cx c	e-cloud 5,	6 mm
Oge .	e-crondag	3 mm
N,	# SLICES	1024
Nm	# MACROP./BUNCH	10240
5×,53	CHROMATICITY	(VARIABLE)
me	E CLOUD DENSIT	(VARIABLE)



PEP-II ECE SIMULATIONS (YUNHAI CAI)

Fix: $M_e = 8 \times 10^5 \text{ cm}^{-3}$, $\xi_x = 0$

(NOMINAL: 500) = 0.5 mm, 500 = 0.14 mm)



CONCLUSIONS

- CLEAR THRESHOLD BEHAVIOR OF of ms. me AT me ~ 5×10° e/cc
- CLEAR BENEFICIAL EFFECT FROM \$ >0
- RESULTS ARE PRELIMINARY
- NEED TO A BETTER DETERMINATION OF Me

(TAKE INTO ACCOUNT VARIOUS ELEMENTS ALONG THE RING)

- RECONCILE WITH OBSERVATIONS

ECE IN LHC: POWER DEPOSITION

F. ZIMMERMANN

O. BRÜNING

G. RUMOLO

H. FURHAN

M. PIVI

R= PHOTON REFLECTI VITY

POWER DEP. IN LHC ARCS

		A		(N
R	860)	Ŝ	DRIFT	BEND
0.1	0	1.9	5.2	2.3
0.1	b	_ 1.1 _	0.67	0.03
0.1	0.6	1.9	14.7	5.44
0.1	0.6	1.1	1.9	0.3
0.05	0.6	1.9	14.7	3.54
0.05	0.6	1.1	1.9	0.22

(F.ZIMMERMANN, CHAMONIX 2004)

ECE AT PEP-II + KEK-B

EN & BEAM BLOWUP (SINGLE-BUNCH EFFECT)

FROM & SUCKED INTO THE BUNCH

SIMULATIONS: Y.CAI

F.ZIMMERMANN } PRL 2000 K. OHMI

~ HEAD-TAIL EFFECT

PERTE:

*ALSO POSSIBLE: COHERENT MULTIBUNCH INSTABILITY (+CONTROLLED BY FEEDBACK)

* NONLINEAR PRESSURE RISE
- DEPENDS ON: LOCATION

BUNCH CURRENT
FILL PATTERN

IN CONCLUSION ...

- 1) AVOID BIM (ADJUST SB, N) IF SEFF >1
- 2) CHOOSE MATERIALS / COATINGS WITH LOW SEY
- 3) " " " Y

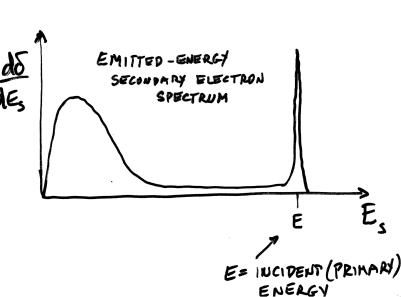
 (Y = PHOTOELECTAL YIELD)
- 4) IF SYNCHROTRON RADIATION HAS ECRIT > 4 eV → USE ANTECHAMBER
- 5) PHOTON REFLECTION COEFF. R MAY BE IMPORTANT

 HIC 25M3
- GAP) -> OTHER DETAILS OF SEY BECOME

 PSR, 100M

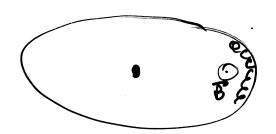
IMPORTANT :

Es = SEi



REMEDIES

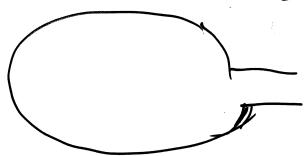
1) WEAK SOLENOIDAL FIELD (PEP-II, KEK-B)
B-20-30 G



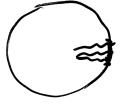
WORKS "VERY WELL"

(TRAP ELECTRONS NEAR WALLS)

2) LOW-SEY COATING (PEP-II) (TIN)
(IF ALUMINUM)



- 3) ANTECHAMBER (PEP-II) AM EXTRACT ~99% OF SR. PHOTONS
- 4) SAWTOOTH SURFACE (LHC, KER-B(?))



1 100h

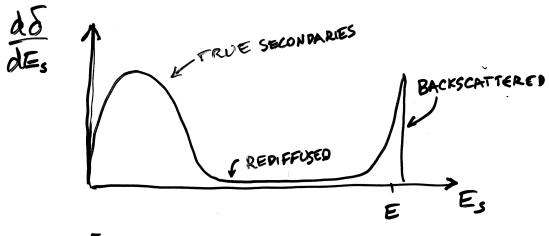
(DECREASE R + Y

IMPORTANT RESEARCH NEEDED:

- MEASUREMENTS OF SEY S(E),

ESPECIALLY &, E , AND & FOR E & 10 eV

- MEASUREMENTS OF EMITTED-ENERGY SPECTAUM



$$\delta(E) = \int_{0}^{E} dE, \frac{d\delta}{dE}$$

FURTHER

- CALIBRATE SIMULATION CODES US. MEASUREMENT

ESP. POWER DEPOSITION (SPS PROGRAM)